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This paper presents results of one phase of research carried out at the Jet Propulsion Laboratory, California Institute of Technology, under Contract No. NAS 7-100, sponsored by the National Aeronautics and Space Administration.

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June 1963

## Photolysis of Nitrogen Dioxide at 3660 and 4047 Å at 25°C\*

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Nitrogen dioxide was irradiated at 3660 and 4047 Å at various pressures of NO<sub>2</sub>, with and without CO<sub>2</sub>, NO, and N<sub>2</sub> as added gases. The data indicate a primary dissociation at 3660 Å and shorter wavelengths corresponding to energies greater than the ON-O bond energy. At 4047 Å, the data are explained by an excited-molecule mechanism. Isotopic oxygen scrambling experiments at 4047 Å indicate the probability of reactions to produce oxygen atoms at that wavelength, but photolysis of NO<sub>2</sub> at trace concentrations and high inert-gas pressures indicates that these atoms are not derived from the unimolecular decomposition of the photoactivated molecule. The inhibition of the quantum yield by NO addition was pressure-dependent, lending further support to the premise that oxygen atoms are important in the mechanism at 4047 Å. It is postulated that the reaction NO<sub>2</sub>\* + NO<sub>2</sub> → N<sub>2</sub>O<sub>3</sub> + O is the source of atomic oxygen.

### INTRODUCTION

SEVERAL investigators have studied the photolysis of NO<sub>2</sub> under various conditions,<sup>1-9</sup> and a fairly consistent set of data has been evolved. Ford<sup>8</sup> has summarized most of the recent work, and has proposed mechanisms for the photolysis between 3100 and 3700 Å. The present study and that of Blacet, Hall, and Leighton<sup>9</sup> essentially substantiate Ford's mechanism in which NO<sub>2</sub> dissociates into NO and O in the primary process at wavelengths which correspond to the bond energy or greater. For the case where NO and O<sub>2</sub> are negligible, O atoms subsequently react with NO<sub>2</sub> which dissociates further via an activated NO<sub>3</sub> molecule as an intermediate. This mechanism is discussed in detail further on.

At wavelengths greater than about 3944 Å, which corresponds to 72.5 kcal/mole which is ΔE for the ON-O bond,<sup>10,11</sup> the data can be explained by an activated NO<sub>2</sub> molecule mechanism. Hall<sup>4</sup> reported that O<sup>18</sup>-O<sup>18</sup> was scrambled at 3100 Å to support the primary dissociation process at that wavelength; how-

ever, he did not observe scrambling at 4047 Å. Further work in this area was thought to be necessary in order to help clarify the mechanism at 4047 Å.

The present work was undertaken to obtain a systematic set of data at both 3660 and 4047 Å to help resolve the rather complex mechanisms that have been proposed. A further purpose of this study was to carry out experiments in the presence of nitrogen, since the results with this gas are directly related to the problems arising from nitrogen dioxide photochemistry in the atmosphere.

### EXPERIMENTAL

The photolysis reactions were carried out in a quartz cell which was 10 cm long and 4 cm in diameter with a volume of 128 ml. The temperature of the system was controlled at 25±0.1°C by means of a thermostat from which water was circulated around the cell.

Quantities of NO<sub>2</sub>, N<sub>2</sub>, NO, and CO<sub>2</sub> were introduced into the cell by standard high-vacuum techniques. The vacuum system was free of mercury since diffusion-pump oil and Kel-F grease were used. The NO<sub>2</sub> pressure was measured by means of a silicone oil manometer (Dow Corning 704, ρ=1.066 g/cc) and also by its optical density at 4350 Å. The cell was designed to fit directly into a Cary Model 11 recording spectrophotometer. The extinction coefficient for NO<sub>2</sub> at 4350 Å was determined to be 1.73×10<sup>5</sup> ml mole<sup>-1</sup> cm<sup>-1</sup>, which is in agreement with the value reported by Hall.<sup>4</sup>

The cell was irradiated with light from a Bausch and Lomb monochromator (f/4.4, dispersion 16 Å/mm), with a General Electric BH6 mercury lamp as the light source. A slitwidth of 1.00 mm was used at both the entrance and exit slits. The total intensity of the beam emerging from the monochromator was measured with a calibrated Eppley line thermopile. The incident light was assumed to be proportional to the microvolt reading on a recording potentiometer. The absolute value of the light intensity was deter-

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<sup>1</sup> H. H. Holmes and F. Daniels, J. Am. Chem. Soc. **56**, 630 (1934).

<sup>2</sup> R. G. Dickinson and W. P. Baxter, J. Am. Chem. Soc. **50**, 774 (1928).

<sup>3</sup> R. G. W. Norrish, J. Chem. Soc. **51**, 1158 (1929).

<sup>4</sup> T. C. Hall, Jr., Ph.D. dissertation, University of California, Los Angeles (October 1953).

<sup>5</sup> H. W. Ford, G. J. Doyle, and N. Endow, J. Chem. Phys. **26**, 1337 (1957).

<sup>6</sup> H. W. Ford and N. Endow, J. Chem. Phys. **27**, 1156 (1957).

<sup>7</sup> H. W. Ford and N. Endow, J. Chem. Phys. **27**, 1277 (1957).

<sup>8</sup> H. W. Ford, Can. J. Chem. **38**, 1780 (1960).

<sup>9</sup> F. E. Blacet, T. C. Hall, and P. A. Leighton, J. Am. Chem. Soc. **84**, 4011 (1962).

<sup>10</sup> F. D. Rossini, D. D. Wagman, W. H. Evans, S. Levine, and I. Jaffe, "Selected Values of Chemical Thermodynamic Properties", Nat. Bur. Std. (U.S.), Circ. 500.

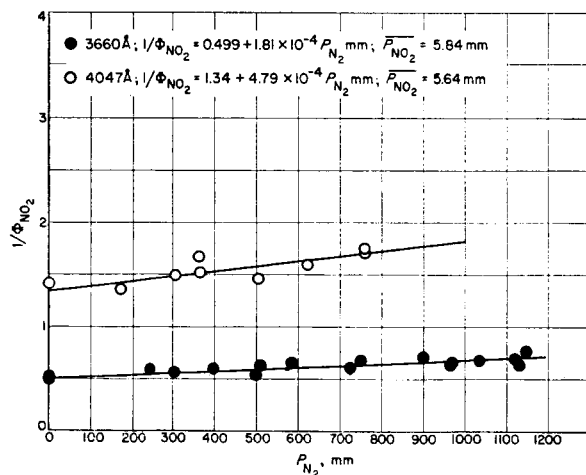
<sup>11</sup> P. Brix and G. Herzberg, J. Chem. Phys. **31**, 1371 (1959).

TABLE I. Least-squares linear relationships in NO<sub>2</sub> photolysis.

Wave-length (Å)	Added gases	Equation of line ( $P$ in mm)	Average $P_{\text{NO}_2}$ (mm)
3660	N <sub>2</sub>	$1/\Phi_{\text{NO}_2} = 0.499 + 1.81 \times 10^{-4} P_{\text{N}_2}$	5.84
3660	CO <sub>2</sub>	$1/\Phi_{\text{NO}_2} = 0.527 + 1.76 \times 10^{-4} P_{\text{CO}_2}$	6.12
3660	None	$1/\Phi_{\text{NO}_2} = 0.52 + 0.233 \text{ } 1/P_{\text{NO}_2}$	
3660	N <sub>2</sub> (const. 1 atm)	$1/\Phi_{\text{NO}_2} = 0.63 + 0.395 \text{ } 1/P_{\text{NO}_2}$	
3660	NO (N <sub>2</sub> const. 1 atm)	$1/\Phi_{\text{NO}_2} = 0.61 + 0.326 P_{\text{NO}}$	6.28
3660	NO (N <sub>2</sub> const. 0.1 atm)	$1/\Phi_{\text{NO}_2} = 0.45 + 0.112 P_{\text{NO}}$	6.01
4047	None	$1/\Phi_{\text{NO}_2} = 1.28 + 0.260 \text{ } 1/P_{\text{NO}_2}$	
4047	N <sub>2</sub> (const. 1 atm)	$1/\Phi_{\text{NO}_2} = 1.32 + 0.741 \text{ } 1/P_{\text{NO}_2}$	
4047	CO <sub>2</sub> (const. 1 atm)	$1/\Phi_{\text{NO}_2} = 1.46 + 0.818 \text{ } 1/P_{\text{NO}_2}$	
4047	N <sub>2</sub>	$1/\Phi_{\text{NO}_2} = 1.35 + 3.60 \times 10^{-4} P_{\text{N}_2}$	5.69
4047	CO <sub>2</sub>	$1/\Phi_{\text{NO}_2} = 1.37 + 6.97 \times 10^{-4} P_{\text{CO}_2}$	8.81
4047	NO (N <sub>2</sub> const. 1 atm)	$1/\Phi_{\text{NO}_2} = 1.22 + 1.97 P_{\text{NO}}$	5.90
4047	NO (N <sub>2</sub> const. 0.1 atm)	$1/\Phi_{\text{NO}_2} = 1.07 + 0.558 P_{\text{NO}}$	6.24

mined by a standard uranyl oxalate actinometric procedure.<sup>12</sup>

The NO<sub>2</sub> used in this work was Matheson Company research grade, and was purified by the method of Harris *et al.*<sup>13</sup> The purified NO<sub>2</sub> was stored as a white

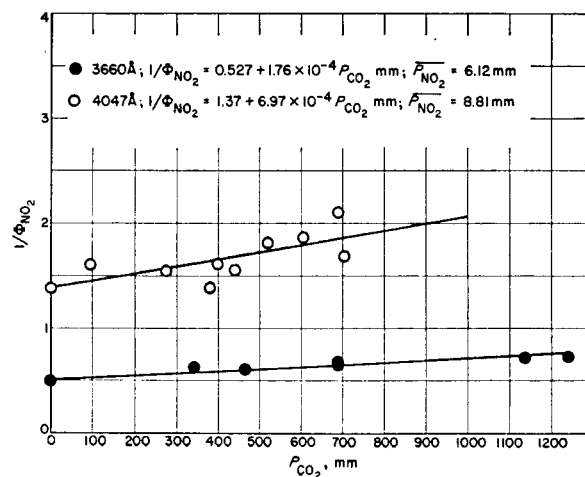
FIG. 1. Dependence of  $1/\Phi_{\text{NO}_2}$  on the nitrogen pressure.

<sup>12</sup> W. A. Noyes, Jr. and P. A. Leighton, *The Photochemistry of Gases* (Reinhold Publishing Corporation, New York, 1941), pp. 82-83.

<sup>13</sup> L. Harris, G. W. King, W. S. Benedict, and R. W. B. Pearse, *J. Chem. Phys.* 8, 765 (1940).

solid at  $-195^\circ\text{C}$  and light was excluded from the storage bulb. Linde high-purity dry nitrogen was passed through ascarite and further dried by passage through CaSO<sub>4</sub> and Mg(ClO<sub>4</sub>)<sub>2</sub>. In the initial runs, the nitrogen was passed through electrolytic Cu at  $800^\circ\text{C}$  to remove possible O<sub>2</sub>, and the resulting N<sub>2</sub> was distilled. Mass-spectrographic analysis showed that the latter purification steps were unnecessary, since impurities such as H<sub>2</sub>O, O<sub>2</sub>, and hydrocarbons were below the limit of detection by ordinary mass-spectrographic techniques (10-50 ppm). The CO<sub>2</sub> was obtained from the National Cylinder Gas Company and was approximately 99.5% pure. It was passed through CaSO<sub>4</sub>, solidified, and pumped on. The middle portion of the sublimed sample was used in the experiments.

NO, obtained from Matheson, was purified by passing the gas through Ascarite, CaSO<sub>4</sub>, Mg(ClO<sub>4</sub>)<sub>2</sub>, and

FIG. 2. Dependence of  $1/\Phi_{\text{NO}_2}$  on the carbon dioxide pressure.

a series of dry-ice-acetone traps. The remaining NO was distilled several times from liquid-nitrogen traps, each time retaining only the middle fractions. The resulting NO was stored as a gas. It was bluish white as a solid and blue as a liquid.

The isotopic oxygen used in the scrambling experiments was obtained from the Bio-Rad Laboratories (32nd and Griffin, Richmond, California). It contained 98.28% O<sup>18</sup>, 0.47% O<sup>17</sup>, and the balance O<sup>16</sup>.

In a typical run, the desired gases were introduced into the sample cell while it was maintained at  $25^\circ\text{C}$ , and the concentration of the NO<sub>2</sub> was measured by means of the Cary spectrometer before and after the experiments. The incident-light intensity from the monochromator was measured by the thermopile output and recorded. The time of exposure was measured by means of a stopwatch which was calibrated in 0.1-sec increments. Quantum yields were calculated by the standard method, corrections for the dissociation of N<sub>2</sub>O<sub>4</sub> and N<sub>2</sub>O<sub>3</sub> being made where necessary. Correction was also made for the absorption of light

TABLE II. Results of oxygen scrambling experiments.

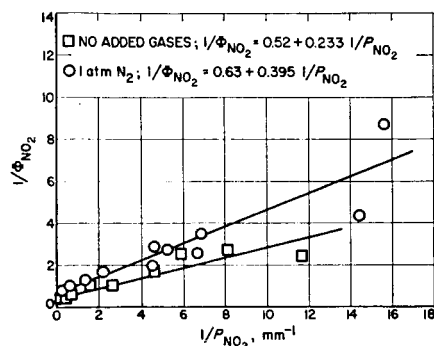
Wavelength Å	$P_{\text{NO}_2}$ (mm)	$P_{\text{O}_2}$ (mm)	$P_{\text{NO}}$ (mm)	Before ratio 36/34	After ratio 36/34	$\Phi_{\text{NO}_2}$	$\Phi_{\text{O}_2^{36}}$	Moles photo- produced $\text{O}_2$	Quanta absorbed/ min	Time in min
4047	8.47	0.71	...	25.0	23.56	0.70	0.062	$1.74 \times 10^{-6}$	$7.49 \times 10^{16}$	40
4047	9.05	0.55	...	25.0	9.77	0.65	0.021	$6.27 \times 10^{-6}$	$8.47 \times 10^{16}$	138
4047	8.41	0.71	...	25.0	8.24	0.73	0.037	$6.16 \times 10^{-6}$	$8.00 \times 10^{16}$	127
4047	8.50	0.67	...	20.8	15.18	0.78	0.035	$3.07 \times 10^{-6}$	$7.92 \times 10^{16}$	60
4358	8.88	0.39	...	20.6	20.8	0	0	0	0	60
4358	8.98	0.28	...	20.6	20.6	0	0	0	0	56
Dark	4.96	0.47	0.20	40.6	35.8	...	...	...	...	57
Dark	4.98	0.51	0.18	40.6	34.0	...	...	...	...	45

by  $\text{N}_2\text{O}_4$  at 3660 Å using Hall's<sup>4</sup> extinction coefficients for  $\text{NO}_2$  and  $\text{N}_2\text{O}_4$ .

The oxygen-scrambling experiments were carried out by introducing about 0.5 mm of the  $\text{O}_2^{36}$ -rich oxygen into the photolysis cell by means of a special vacuum-tight hypodermic syringe. The samples were irradiated as usual, and quantum yields were determined as described above. At the end of a run,  $\text{O}_2$  was separated from the  $\text{NO}_2$  and  $\text{NO}$  by freezing out the oxides in a liquid-nitrogen trap. The ratios of the  $\text{O}_2^{36}$  to  $\text{O}_2^{34}$  peaks were determined by ordinary mass-spectrographic analytical techniques. The values of the ratios after photolysis were compared with those before photolysis to estimate the extent of scrambling.

### RESULTS

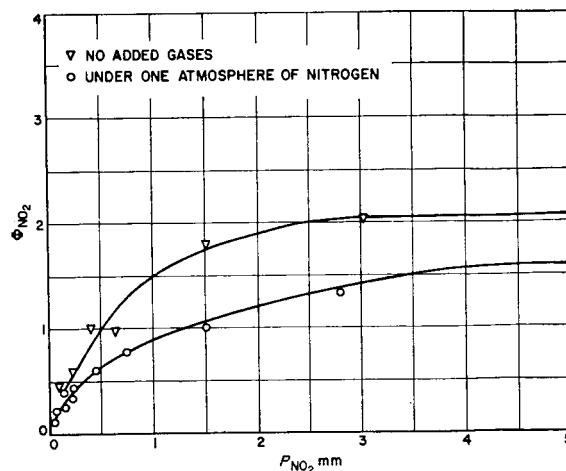
The quantum yield for  $\text{NO}_2$  dissociation was measured at 3660 Å and at constant  $\text{NO}_2$  pressure ( $\sim 6$  mm) as a function of  $\text{CO}_2$  and  $\text{N}_2$  pressure. The results are given in Table I and in Figs. 1 and 2. The data suggest linear relationships which are substantially in agreement with the pressure dependence reported by Blacet, Hall, and Leighton<sup>9</sup> for the effect of  $\text{CO}_2$  at 3100 Å. In addition, the data indicate that the third-body efficiency of nitrogen is approximately the same as for  $\text{CO}_2$  under these conditions.

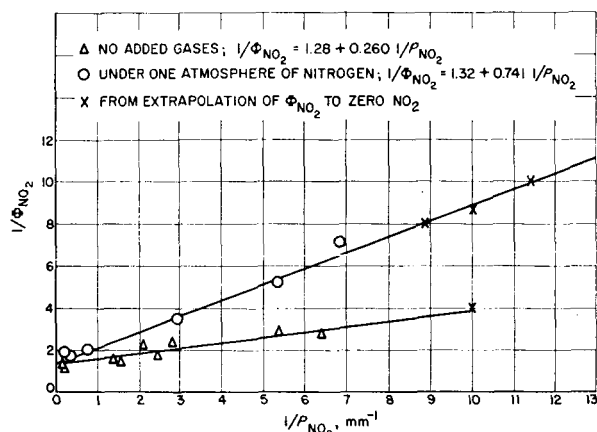
FIG. 3. Dependence of  $1/\Phi_{\text{NO}_2}$  on  $1/P_{\text{NO}_2}$  at 3660 Å.

Quantum yield measurements were also made with and without  $\text{N}_2$ , varying the  $\text{NO}_2$  pressure. These results are shown in Table I and in Figs. 3 and 4. A plot of  $1/\Phi_{\text{NO}_2}$  against  $1/(\text{NO}_2)$  yields a straight line. However, the data at low pressures of  $\text{NO}_2$  are subject to rather large errors because of the small concentration changes which are being measured. Therefore, Fig. 4 indicates a trend which must be verified by more precise measurements at pressures below about 0.1 mm of  $\text{NO}_2$ . The trend in the quantum yield toward zero as  $(\text{NO}_2)$  approaches zero is not anticipated by the present mechanism or by other experimental data.<sup>6</sup>

The mechanism of the photolytic dissociation of  $\text{NO}_2$  at 4047 Å is not as clearly understood as that at 3660 Å and shorter wavelengths. Therefore, quantum yield measurements were made under a variety of conditions to help determine the mechanism at 4047 Å, and in particular, to attempt to distinguish between the postulated activated molecule<sup>9</sup> and the atomic-oxygen mechanism which obtains at shorter wavelengths.

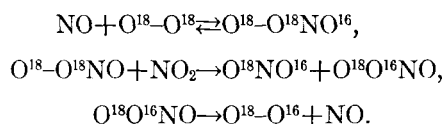
Accordingly, a series of oxygen-scrambling experiments were performed using the change in the ratio of

FIG. 4. Dependence of  $\Phi_{\text{NO}_2}$  on the  $\text{NO}_2$  pressure at 3660 Å.

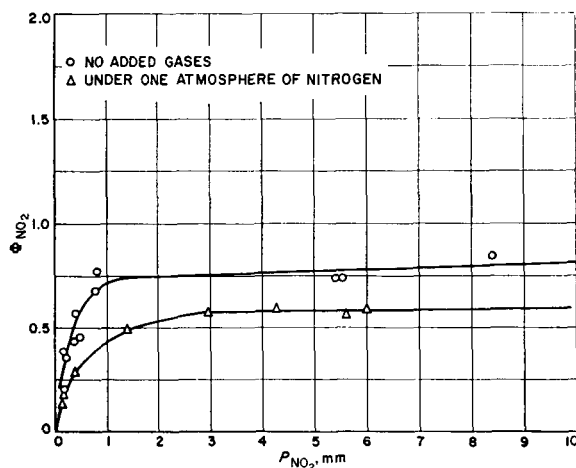
FIG. 5. Dependence of  $1/\Phi_{\text{NO}_2}$  on  $1/P_{\text{NO}_2}$  at 4047 Å.

$\text{O}_2^{36}$  to  $\text{O}_2^{34}$ , as determined by mass spectrographic analysis, to follow the scrambling. The results of these experiments are summarized in Table II. The data show that scrambling is appreciable as long as there is a quantum yield for  $\text{NO}_2$  dissociation. However, when the quantum yield approaches zero at 4358 Å, no measurable scrambling occurs. It is, therefore, evident that the scrambling phenomena are directly a result of the mechanism of photolysis.

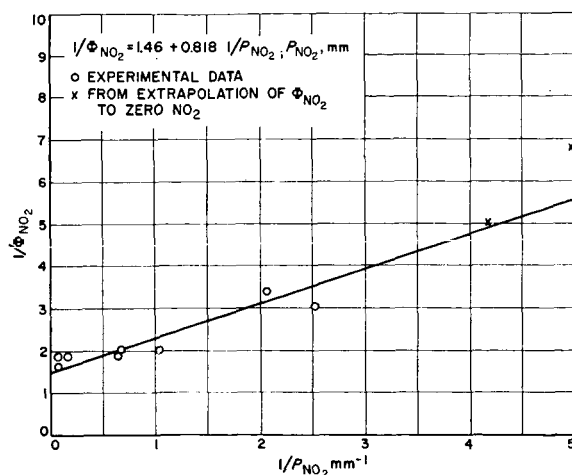
The work of R. A. Ogg, Jr.,<sup>14</sup> suggests that scrambling can take place in the dark in mixtures of NO,  $\text{NO}_2$ , and  $\text{O}_2^{36}$ . The proposed mechanism for scrambling was



Therefore, two runs were made in the dark under

FIG. 6. Dependence of  $\Phi_{\text{NO}_2}$  on the  $\text{NO}_2$  pressure at 4047 Å.

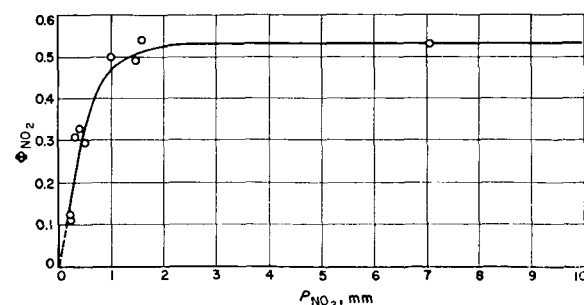
<sup>14</sup> R. A. Ogg, Jr., J. Chem. Phys. **21**, 2079 (1953).

FIG. 7. Dependence of  $1/\Phi_{\text{NO}_2}$  on  $1/P_{\text{NO}_2}$  at 4047 Å and under 1 atm of carbon dioxide.

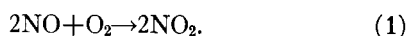
conditions simulating those in which photolysis took place. The NO concentration was approximately equal to the average which would be photoproduced during 1 h. Mass-spectrographic analysis showed that about  $1 \times 10^{-7}$  moles/liter of  $\text{O}_2^{36}$  were lost in each case during 1-h periods. However, in a comparable period,  $3 \times 10^{-6}$  moles/liter of  $\text{O}_2^{36}$  were lost during photolysis. This result indicates that oxygen atoms are probably responsible for most of the scrambling that was observed and that the dark reactions do not predominate.

Several series of experiments were carried out at 4047 Å to further determine the mechanism of photolysis.  $\text{NO}_2$  was photolyzed at constant  $\text{NO}_2$  pressure ( $\sim 6$  mm) while the  $\text{N}_2$  and  $\text{CO}_2$  pressures were varied from zero to about 1 atm. The data yield straight lines, the equations for which are given in Table I and shown in Figs. 1 and 2. In addition, experiments were performed with variable  $\text{NO}_2$  pressure without any added gases and with  $\text{N}_2$  and  $\text{CO}_2$  each constant at one atmosphere pressure. Plots of  $1/\Phi_{\text{NO}_2}$  as a function of  $1/(\text{NO}_2)$  yielded straight lines. These data are shown in Table I and in Figs. 5 through 8.

A final series of experiments was carried out at 3660 and at 4047 Å with the addition of NO at partial

FIG. 8. Dependence of  $\Phi_{\text{NO}_2}$  on the  $\text{NO}_2$  pressure at 4047 Å and under 1 atm of carbon dioxide.

pressures low enough to avoid the back reaction:

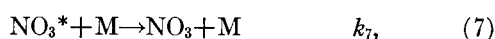
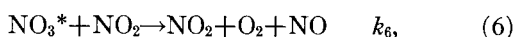


Photolysis with an atmosphere of nitrogen showed a greater inhibition of the quantum yield than that with 0.1 atm. In addition, the relative effects were the same at both wavelengths. This indicated that the influence of NO is pressure-dependent and could further support an oxygen atom mechanism at 4047 Å. The results of these experiments are also given in Table I and in Fig. 9.

## DISCUSSION

### A. Photolysis at 3660 Å

The mechanism of the photolytic dissociation of NO<sub>2</sub> at 3660 Å has been described by Ford<sup>5-8</sup> and reaffirmed by Blacet, Hall, and Leighton.<sup>9</sup> The mechanism may be written as follows:



Assuming steady state for O and NO<sub>3</sub><sup>\*</sup>, the quantum-yield expression is:

$$\Phi_{\text{NO}_2}^{-1} = \frac{1}{2} + \frac{k_4 k_7 (\text{M})}{2[k_3 k_5 + k_6(k_3 + k_4)(\text{NO}_2) + k_3 k_7 (\text{M})]} \quad (9)$$

This mechanism indicates that photodissociation takes place at 3660 Å. Photodissociation would be expected since the Einstein unit at 3660 Å is 78.2 kcal

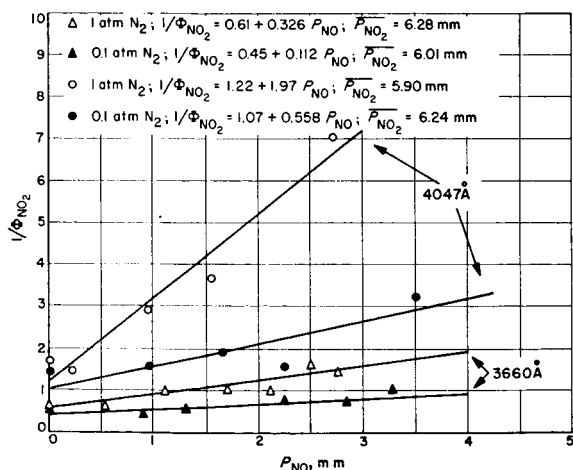


FIG. 9. Dependence of  $1/\Phi_{\text{NO}_2}$  on the NO pressure.

TABLE III. Rate constants for the dissociation of NO<sub>2</sub> at 3660 Å.

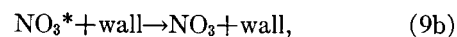
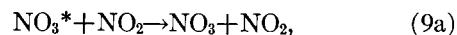
Constant or ratio	Value	Remarks
$k_3$	$2.1 \times 10^9$ liter mole <sup>-1</sup> sec <sup>-1</sup>	Reference 8
$k_4$	$6 \times 10^9$ liter mole <sup>-1</sup> sec <sup>-1</sup>	Reference 8
$k_5/k_6$	0.002 mole liter <sup>-1</sup>	This work
$k_7/k_6$	$7.6 \times 10^{-3}$	From slope in Fig. 1 (M = N <sub>2</sub> ) and constants $k_3$ , $k_4$ , $k_5/k_6$ from this work and references 8 and 9.
$k_6$	$5 \times 10^8$ liter mole <sup>-1</sup> sec <sup>-1</sup>	Reference 8
$k_7$	$3.8 \times 10^6$ liter mole <sup>-1</sup> sec <sup>-1</sup>	From $k_6$ , reference 8 and $k_7/k_6$ this work. (M = N <sub>2</sub> )
$k_8$	$3.4 \times 10^9$ liter mole <sup>-1</sup> sec <sup>-1</sup>	Reference 8

which is 5.7 kcal in excess of the ON-O bond energy. The problem would be settled uniquely if the fate of the O atoms were known. The assumptions that Reactions (3) and (4) take place seems reasonable and are not in conflict with the experimental evidence.

If the value of  $k_7(\text{M})$  is smaller than  $k_5$ , the term  $k_3 k_7 (\text{M})$  may be neglected and Eq. (9) yields a straight line when  $1/\Phi_{\text{NO}_2}$  is plotted as a function of M. Lines of this form are shown in Figs. 1 and 2. The constants given in Table III are taken from earlier estimates and the slope of the 3660-Å line in Fig. 1. The constants derived from this work are in good agreement with those of earlier investigators.<sup>8,9</sup>

The mechanism proposed above fits the data fairly well for NO<sub>2</sub> pressures greater than about 2 mm. However, Eq. (9) requires that, as M approaches zero,  $1/\Phi_{\text{NO}_2}$  approaches 0.5 and is not a function of the NO<sub>2</sub> pressure. Figures 3, 5, and 7 show that  $1/\Phi_{\text{NO}_2}$  is a function of the NO<sub>2</sub> pressure below about 2 mm. This would require an extra term in Eq. (9) that is proportional to  $1/(\text{NO}_2)$ . Blacet, Hall, and Leighton<sup>9</sup> have proposed a wall deactivation to account for the phenomenon. However, since a similar effect is evident with and without added N<sub>2</sub> up to one atmosphere of N<sub>2</sub>, wall reactions may not be the only explanation for the results.

If one adds the following steps to the mechanism,



the quantum yield expression becomes

$$\Phi_{\text{NO}_2}^{-1} = \frac{1}{2} + \frac{k_4 k_{9a} (\text{NO}_2) + k_4 k_7 (\text{M}) + k_4 k_{9b}}{2k_3[(k_6 + k_{9a})(\text{NO}_2) + k_5 + k_{9b} + k_7 (\text{M})] + 2k_4 k_6 (\text{NO}_2)} \quad (9c)$$

This relationship would account for the dependence

of the quantum yield on the  $\text{NO}_2$  pressure when  $M$  is zero because of the term which includes the deactivation, (9b). When  $M$  is not zero, the term in  $k_7(M)$  would be retained and the deactivation term in (9b) would go to zero. Therefore, the quantum yield would be a function of the  $\text{NO}_2$  pressure under both conditions.

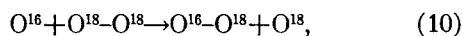
Using the constants in Table III and appropriate data from Table I,  $k_{9a}$  was calculated to be  $5.4 \times 10^7$  liter mole $^{-1}$  sec $^{-1}$  and  $k_{9b}$   $1.6 \times 10^5$  sec $^{-1}$  as calculated from Eq. (9c).

Quantum yield measurements are presently being made at  $\text{NO}_2$  partial pressures well below 2 mm and the results will be reported in the near future.

### B. Photolysis at 4047 Å

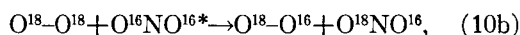
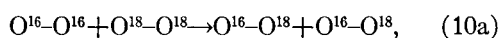
The photolysis of  $\text{NO}_2$  at wavelengths greater than that corresponding to the energy of the ON-O bond (72.5 kcal/mole) has been a subject of great interest since predissociation cannot occur, and activated molecule mechanisms must be proposed. Blacet, Hall, and Leighton,<sup>9</sup> most recently, and Norrish,<sup>3</sup> earlier, have proposed such excited molecule mechanisms.

However, the work of Sato and Cvetanović<sup>15</sup> on the formation of organic oxides during the photolysis of  $\text{NO}_2$  in the presence of olefins suggests that oxygen atoms are important in the reaction. The oxygen scrambling data shown in Table II give qualitative evidence that atoms are present during the photolytic process. The principal scrambling mechanism is probably:



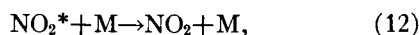
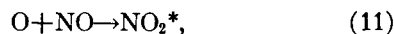
the  $\text{O}^{16}$  being derived from  $\text{NO}_2$ .

Molecular scrambling mechanisms, such as

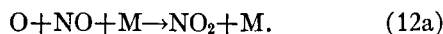


should not occur.

Further evidence for the presence of oxygen atoms is given in Fig. 9. The pressure dependence of the influence of NO on the inhibition of the quantum yield indicates that the following sequence of reactions may occur:



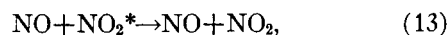
or



When Reaction (12a) is considered along with the mechanism given below, a rate constant,  $12a$ , may be calculated from the data in Fig. 9. This calculation

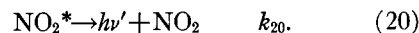
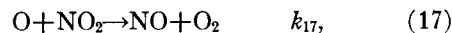
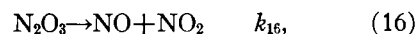
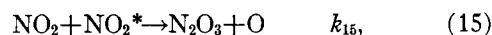
yields approximately  $4 \times 10^{11}$  liter $^2$  mole $^{-2}$  sec $^{-1}$ . This is an order of magnitude greater than the value  $1.8 \times 10^{10}$  liter $^2$  mole $^{-2}$  sec $^{-1}$  reported by Ford<sup>8</sup> and others. Since added NO has a greater effect than added  $\text{N}_2$  when both are present, at least one other process is occurring and  $k_{12a}$  as calculated would be expected to be too high. However, the results support the premise that Reaction (12a) is one cause of the decrease in the quantum yield as a function of NO pressure.

Alternative reactions of NO with  $\text{NO}_2^*$ ,



are not a function of  $M$ , and may be assumed to be no more effective than the deactivation by  $\text{N}_2$  or  $\text{CO}_2$ .

The data may be explained by the following mechanism:



Assuming steady state on O and  $\text{NO}_2^*$ , the resulting quantum yield expression is

$$\Phi_{\text{NO}_2}^{-1} = \frac{1}{2} + \frac{k_{19}}{2k_{16}} + \frac{k_{18}(M)}{2k_{16}(\text{NO}_2)} + \frac{k_{20}}{2k_{15}(\text{NO}_2)}. \quad (21)$$

Reaction (14) is probable since the ON-O bond energy is 72.5 kcal/mole and the energy equivalent to 4047 Å is only 70.7 kcal/mole. Therefore, dissociation would not be expected at this wavelength.

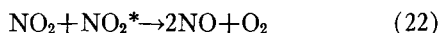
The manner in which dissociation follows the activation of the  $\text{NO}_2$  molecule is, of course, the most impor-

TABLE IV. Rate-constant relationships for  $\text{NO}_2$  photolysis at 4047 Å.

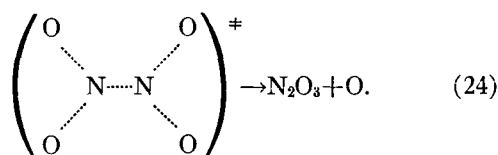
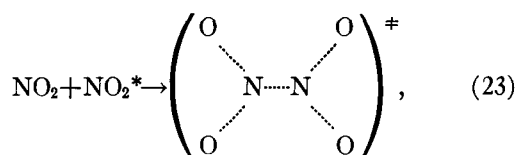
Constant or ratio	Value	Remarks
$k_{17}$ (also $k_3$ )	$2.1 \times 10^9$ liter mole $^{-1}$ sec $^{-1}$	Reference 6
$k_{18}/k_{15}$	$4.1 \times 10^{-3}$	with $M = \text{N}_2$
$k_{18}/k_{16}$	$1.23 \times 10^{-3}$	with $M = \text{CO}_2$
$k_{19}/k_{15}$	1.64	with $M = \text{N}_2$
$k_{19}/k_{16}$	1.92	with $M = \text{CO}_2$
$k_{20}/k_{15}$	$2.8 \times 10^{-5}$ moles liter $^{-1}$	with $M = \text{O}$
$k_{19}/k_{18}$	$4.0 \times 10^3$	with $M = \text{N}_2$
$k_{19}/k_{16}$	$1.56 \times 10^3$	with $M = \text{CO}_2$

<sup>15</sup> S. Sato and R. J. Cvetanović, Can. J. Chem. **36**, 279 (1958).

tant consideration in the mechanism. Blacet, Hall, and Leighton<sup>9</sup> proposed

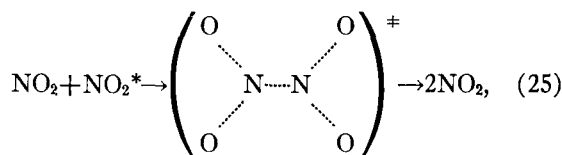


to explain their data. However, this reaction does not explain the presence of oxygen atoms. Furthermore, since the quantum yield falls to zero at 4358 Å, a mechanism is required that is wavelength-dependent such that the energy necessary to break the bond is not available at 4358 Å and longer wavelengths. Reactions (15) and (16) satisfy the two requirements. If one assumes that the reaction proceeds by the formation of a transition complex such as the following:



Approximately 13.3 kcal/mole are gained by the molecules in forming the N-N bond if the energetics of the  $2\text{NO}_2 \rightleftharpoons \text{N}_2\text{O}_4$  reaction may be taken as the model. The dissociation of  $\text{N}_2\text{O}_4$  to form  $\text{N}_2\text{O}_3$  and O requires about 76.7 kcal/mole. This energy may be obtained by adding the N-N bond energy to the energy supplied by the light. At 4047 Å, this gives about 84.0 kcal/mole, which is sufficient to permit reaction. At 4358 Å, the light supplied only 65.7 kcal/mole, yielding a total of 79.0 kcal/mole. If one makes the assumption that more than 2.3 kcal/mole and less than 7.3 kcal/mole in excess of the bond energy are needed for the dissociation, the proposed mechanism may be justified.

The question may be raised whether the competitive reaction,



would be so much faster than that indicated by Reaction (24), that it would predominate. The ratio of rate constants  $k_{19}/k_{15}$  should furnish a partial answer. This ratio is about 1.6 when M is  $\text{N}_2$ , indicating that dissociation may be appreciable relative to deactivation. It is realized that Reaction (25) is only one mode of deactivation which is implied by Reaction (19).

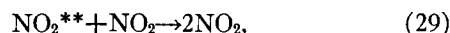
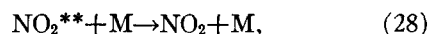
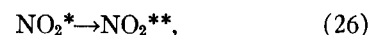
Since the data are not sufficiently sensitive to distinguish the effects of  $\text{NO}_3^*$ , and since the simplified mechanism does explain the data, references to the

formation and subsequent reactions of  $\text{NO}_3^*$  have been omitted from the mechanism above. Inclusion of  $\text{NO}_3^*$  as a long-lived species leads to a complex quantum-yield expression which may or may not fit the data depending on the relative values of certain constants.

The fluorescence step [Reaction (20)] is proposed in accordance with the work of Baxter,<sup>16</sup> who determined the quenching of fluorescence at 4047 Å and at 4358 Å with and without  $\text{H}_2$ ,  $\text{N}_2$ ,  $\text{CO}_2$ , and  $\text{O}_2$  as added gases. Quenching of fluorescence by  $\text{NO}_2$  and by the added gases followed a Stern-Volmer relationship. Appreciable fluorescence was observed with up to 18.3 mm of  $\text{NO}_2$ . With added  $\text{N}_2$ , fluorescence was still observed up to 15.2 mm. These data indicate that fluorescence must be considered as a means of deactivating  $\text{NO}_2^*$  at all of the pressures of  $\text{NO}_2$  used in this work. However, fluorescence may be unimportant at high inert-gas pressures.

Baxter assumed a mean lifetime of fluorescence of  $10^{-7}$  sec. Although this leads to a very large collision diameter,  $30 \times 10^{-8}$  cm, it is not unusual for cases like this. A short lifetime would explain the lack of quenching up to appreciable pressures of  $\text{NO}_2$ .

Steps such as:



were described by Blacet, Hall, and Leighton<sup>9</sup> and are certainly probable here. However, although they are implied, they are not capable of resolution by the present data.

The effect of fluorescence relative to photolysis may be demonstrated by the relative rate constants,  $k_{20}/k_{15}$ . Combination of the expressions  $1/\Phi_{\text{NO}_2}$  vs  $1/(\text{NO}_2)$  at  $\text{M}=\text{O}$  and at one atmosphere of nitrogen gives  $k_{20}/k_{15}=0.52$  mm, which corresponds to  $2.8 \times 10^{-5}$  moles/liter. This is in good agreement with the value of  $3 \times 10^{-5}$  moles/liter given by Blacet, Hall, and Leighton.<sup>9</sup> A summary of rate-constant relationships is given in Table IV.

## CONCLUSIONS

The results of this investigation indicate that the mechanism for the photolysis of  $\text{NO}_2$  at 3660 Å explains the data for pressures of  $\text{NO}_2$  above about 2 mm. Further experiments are necessary to determine the  $\text{NO}_2$  pressure dependence of the quantum yield at pressures below 2 mm; investigating this region will probably show that the mechanism at 3660 Å is more complicated than previously reported.

<sup>16</sup> W. P. Baxter, J. Am. Chem. Soc. **52**, 3920 (1930).



The results at 4047 Å for the photolytic decomposition of NO<sub>2</sub> furnish strong evidence that oxygen atoms are important in the NO<sub>2</sub> dissociation and in the isotopic-oxygen scrambling.

This conclusion is based on the observations of Cvetanović<sup>14</sup> that oxygen atoms are available to form organic oxides when NO<sub>2</sub> is photolyzed in the presence of olefins, on the isotopic oxygen scrambling, on the pressure dependence of the inhibition of the quantum

yield by NO, and by the wavelength dependence of the quantum yield.

#### ACKNOWLEDGMENTS

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